

GRAPPA ANISOTROPY WORKSHOP

Sept. 2013, AMSTERDAM

ANISOTROPIC ORIENTATIONS
OF
POLARISATIONS FROM QUASAR LIGHT

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VINCENT PELGRIMS

in collaboration with J.-R. Cudell

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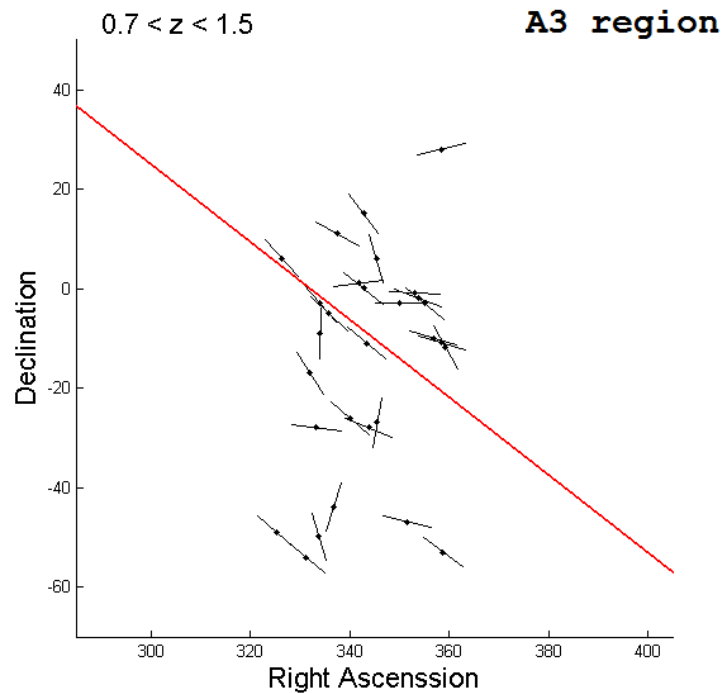


IFPA, AGO Dept.
LIÈGE UNIVERSITY (BELGIUM)



Introduction

The large-scale coherent orientations of quasar polarisation vectors



Probability of uniformity $\sim 6 \cdot 10^{-5}$

[Hutsemékers, D. 1998]

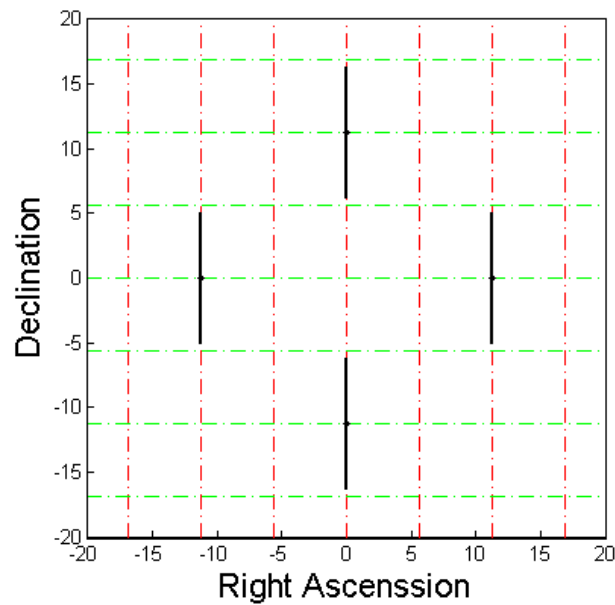
[Hutsemékers, D., Lamy, H. 2001]; [Jain et al. 2004]; [Cabanac et al. 2005]; [Hutsemékers et al. 2005];

- Full sample : 355 quasars
- Orientation correlations up to ~ 1 Gpc at $z \approx 1-2$
- preferred direction \longleftrightarrow redshift slice

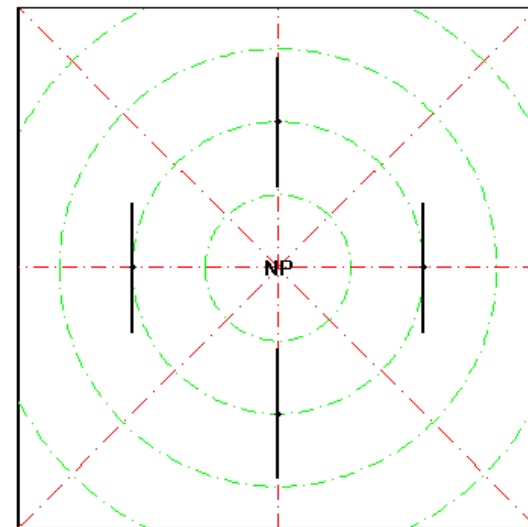
Introduction

Axis dependency

Significance Level \longleftrightarrow coordinate system



$$\psi_{i=1,2,3,4}=0$$



$$\psi_{1,2,3,4}=[0, 90, 0, 90]^\circ$$

A Coordinate-Invariant Method

Need for a **coordinate-invariant statistical test** that keeps **physical information** and that can be used for **non-uniformly dispersed** sources

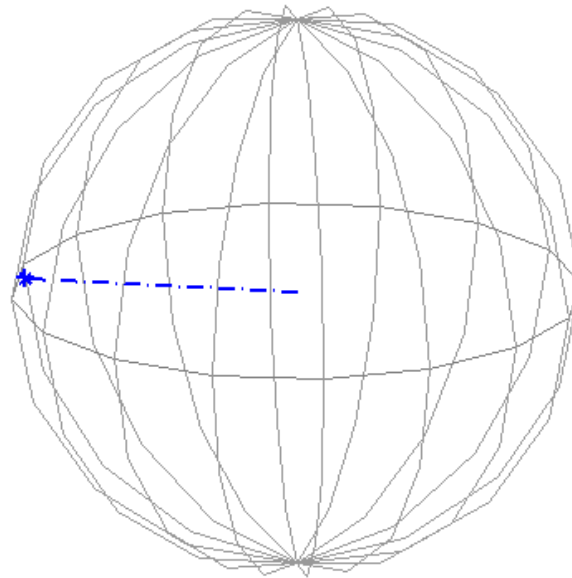
Polarisation Space

Cone Algorithm

The Polarisation Space

Physical information: electric field oscillation direction in **3D**

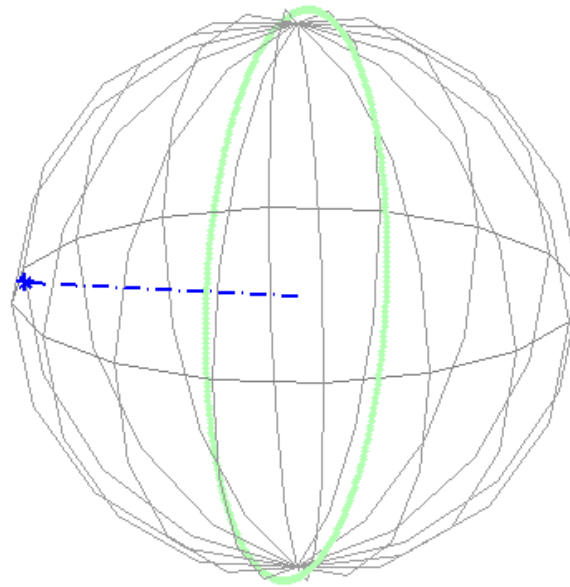
1 data point = 1 line of sight + 1 polarisation angle in the orthogonal plane



The Polarisation Space

Physical information: electric field oscillation direction in **3D**

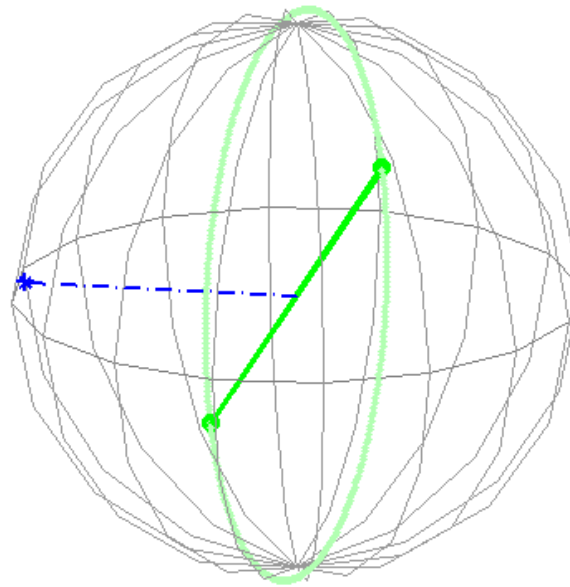
1 data point = 1 line of sight + 1 polarisation angle in the orthogonal plane



The Polarisation Space

The polarisation space = unit 2-sphere

Polarisation “vectors”, regarded as points on the polarisation space



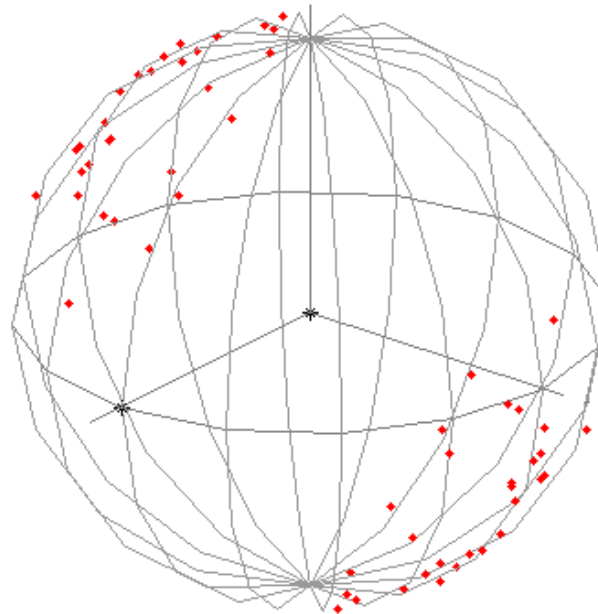
Polarisation axes: circular data ➔ spherical data

The Polarisation Space

Polarisation “vectors”, : spherical data

Configuration \longleftrightarrow density of polarisation points with spherical caps of equal area

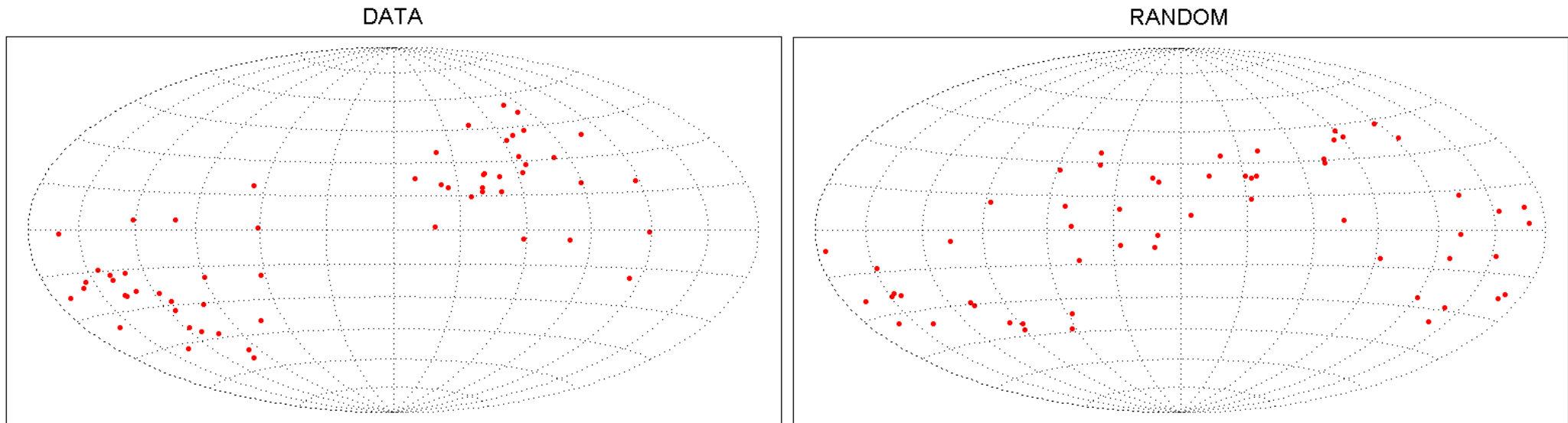
(half aperture angle η)



Polarisation space for the A3 region
from [\[Hutsemékers et al. 2005\]](#)

The Polarisation Space

Polarisation “vectors”, spherical data

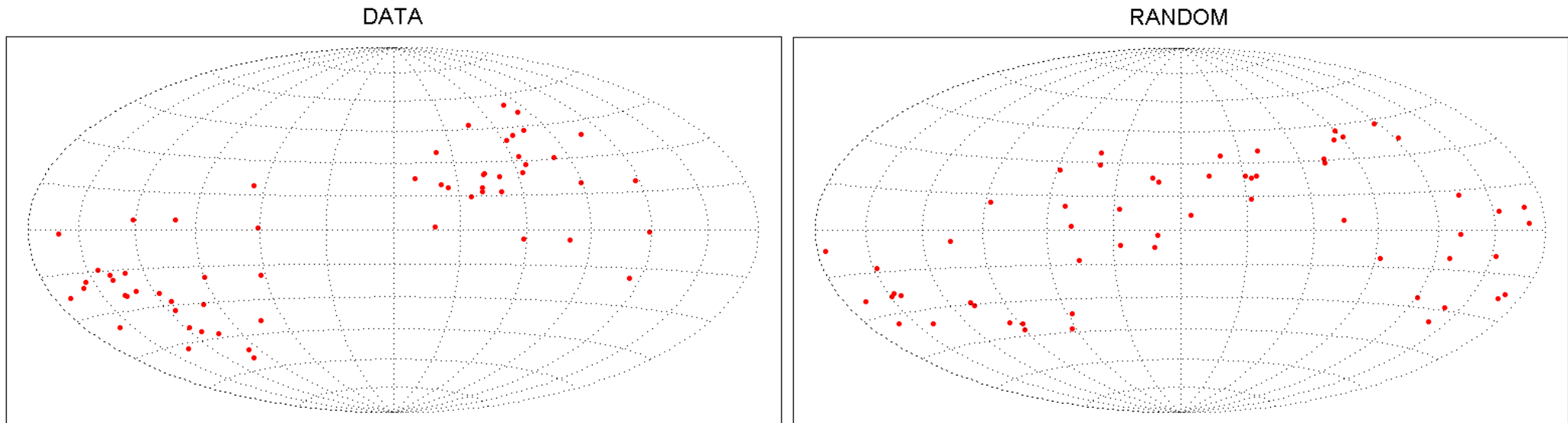


Polarisation space for the **A3 region** from [\[Hutsemékers et al. 2005\]](#)

How to distinguish unexpected densities of polarisation points from “natural,” ones, i.e, drawn from a uniform distribution?

The Polarisation Space

Polarisation “vectors”: spherical data



Polarisation space for the **A3 region** from [\[Hutsemékers et al. 2005\]](#)

How to distinguish unexpected densities of polarisation points from “natural,, ones, i.e, drawn from a uniform distribution?

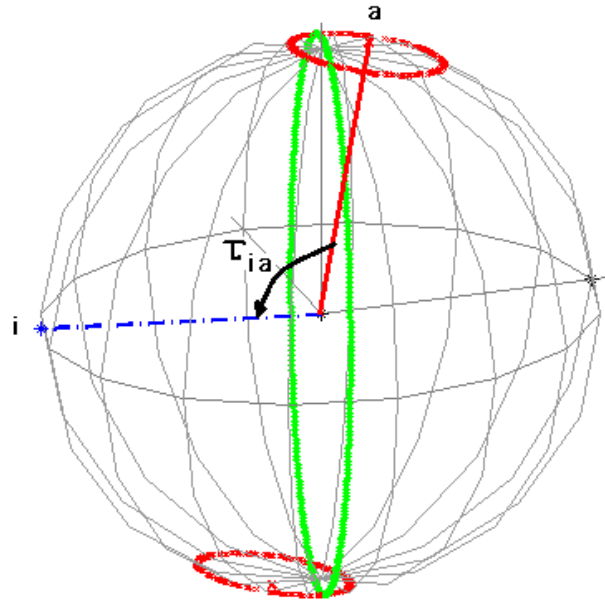


Cone Algorithm

The Cone Algorithm

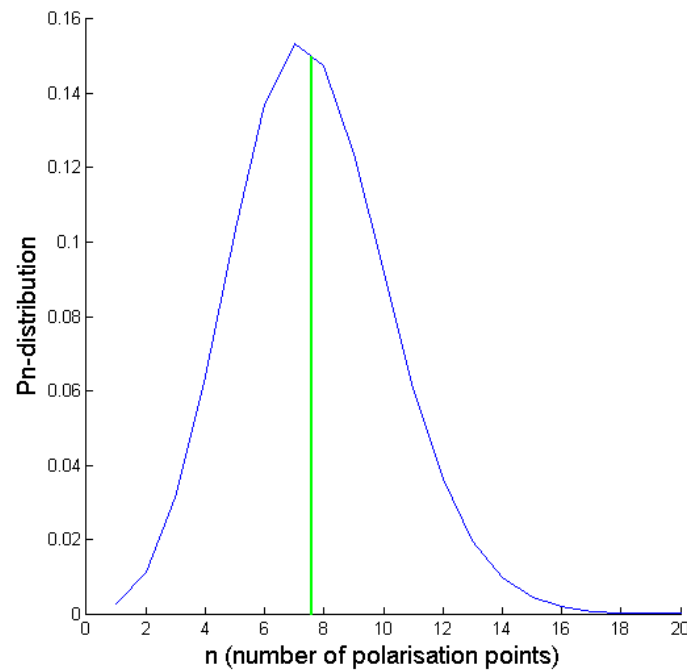
- 1) Cones of fixed half aperture angle η
- 2) Analytical individual probability p_{ia}

$$p_{ia} = \frac{\text{arc-length}_{ia}}{\pi} \propto f(\eta, \tau_{ia})$$



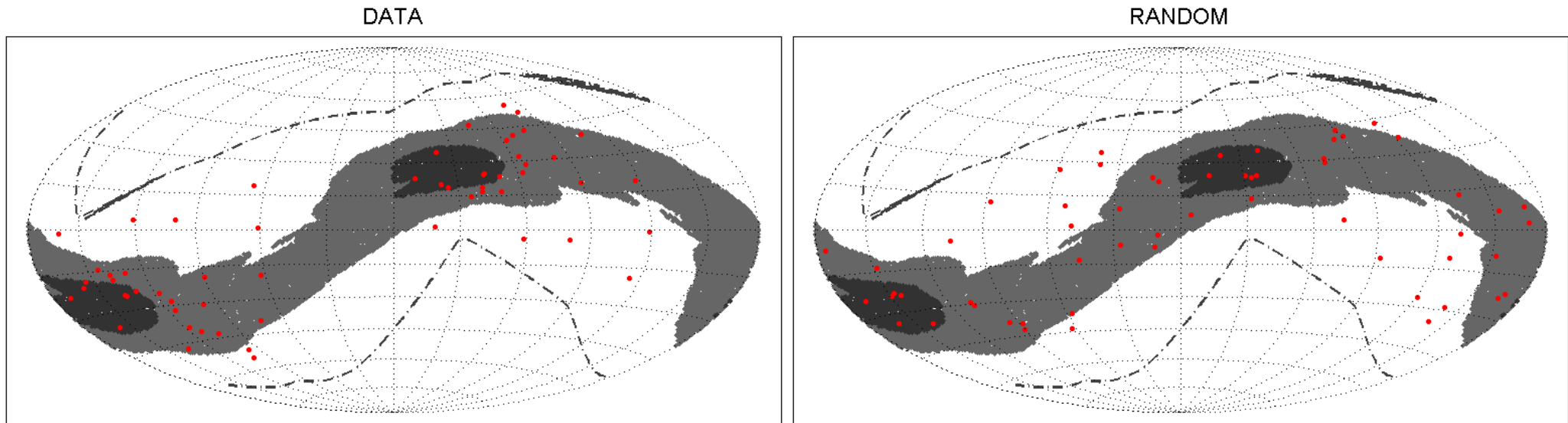
The Cone Algorithm

- 1) Cones of fixed half aperture angle η
- 2) Analytical individual probability p_{ia}
- 3) Recursive algorithm gives probability distribution P_n^a



Polarisation Map

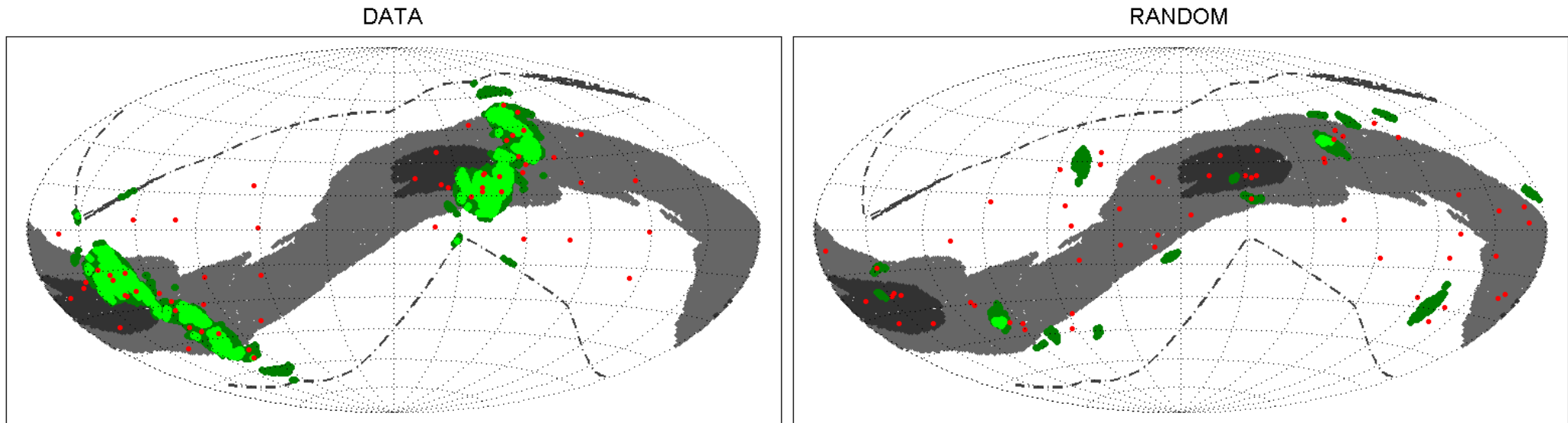
Expected Background



Polarisation space for the **A3 region** from [\[Hutsemékers et al. 2005\]](#)

Polarisation Map

Observations v.s. Expected Background

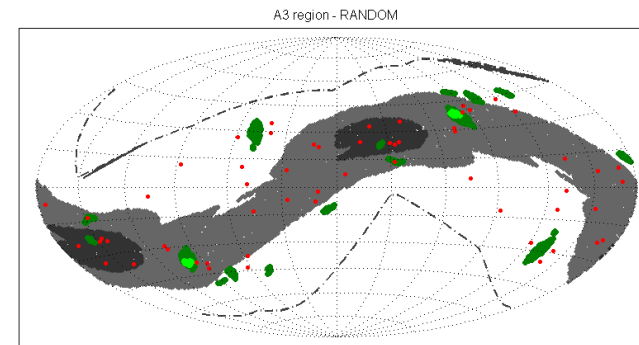
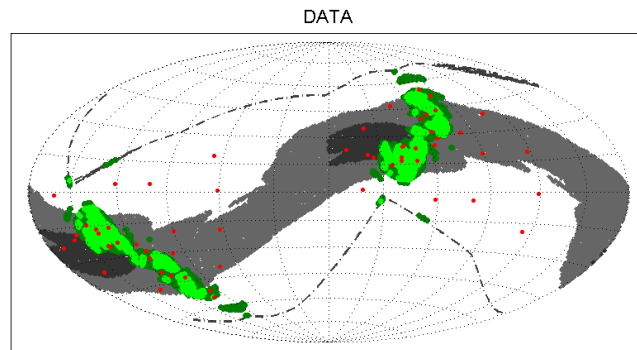


Polarisation space for the **A3 region** from [\[Hutsemékers et al. 2005\]](#)

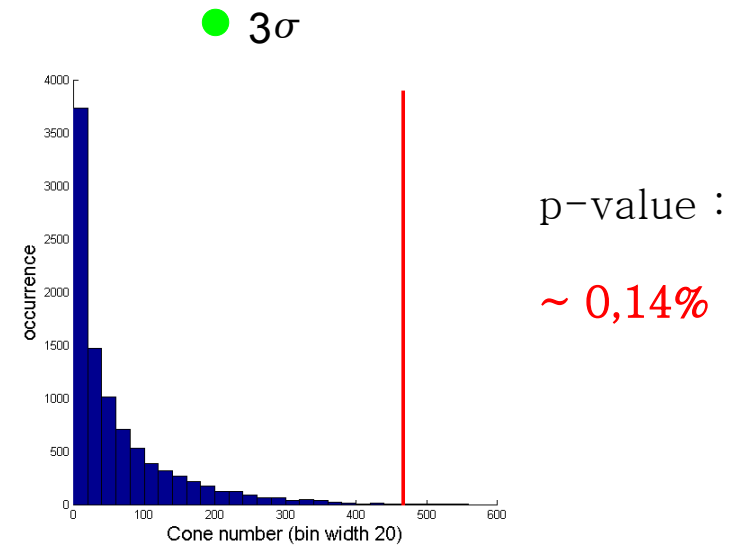
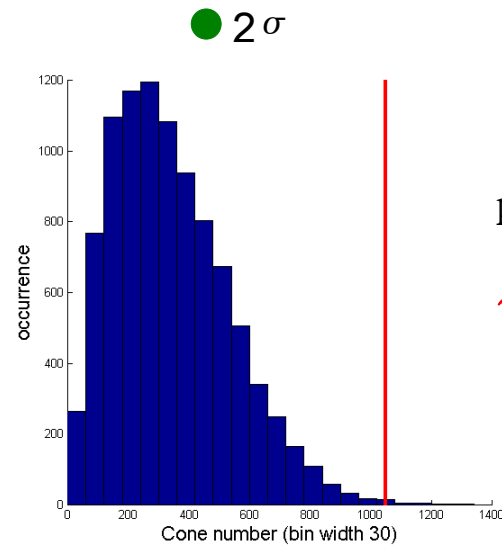
- At least 3σ deviation in number of polarisation points
- At least 2σ deviation in number of polarisation points

Polarisation Map

Observations v.s. Expected Background

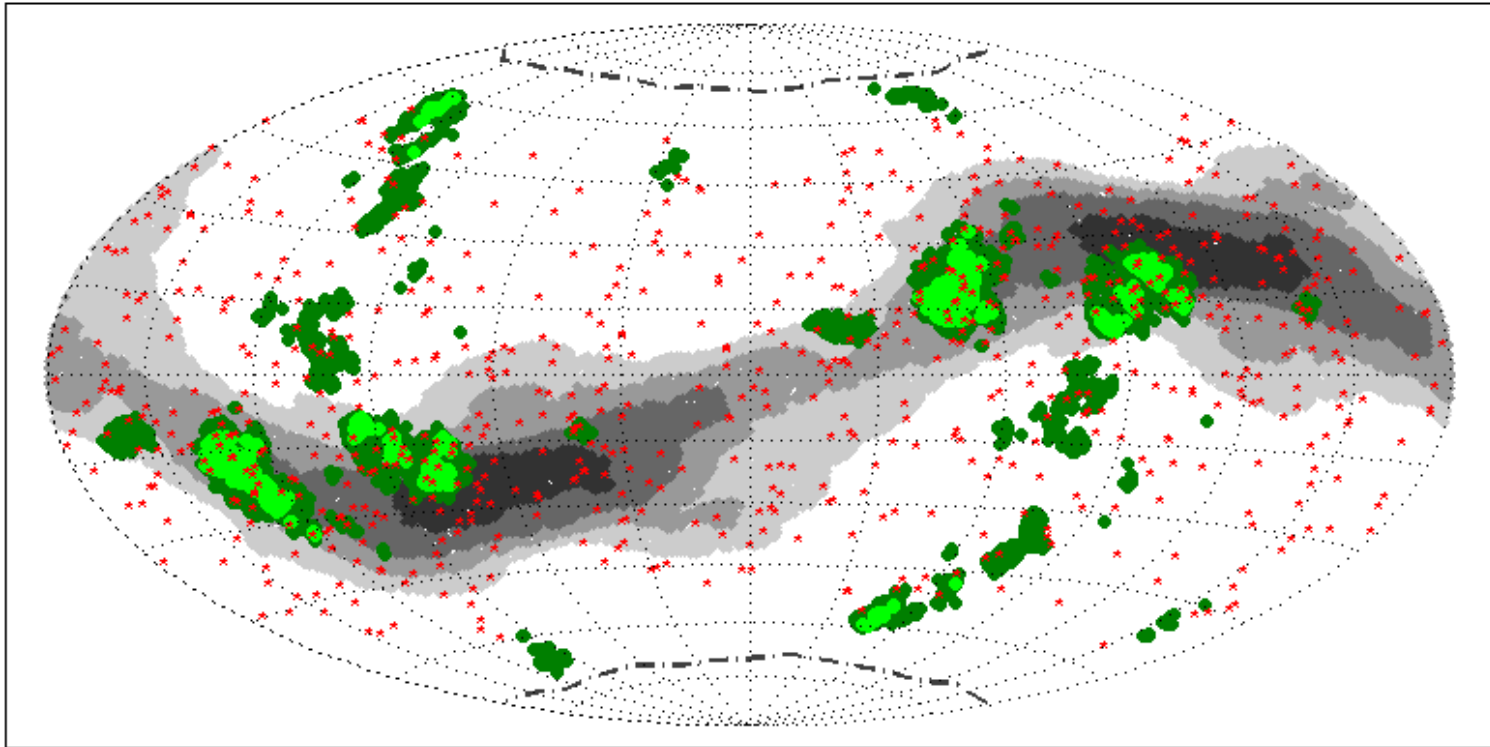


Polarisation space for the A3 region from [\[Hutsemékers et al. 2005\]](#)



Polarisation Map Whole Sample

(355 quasars)

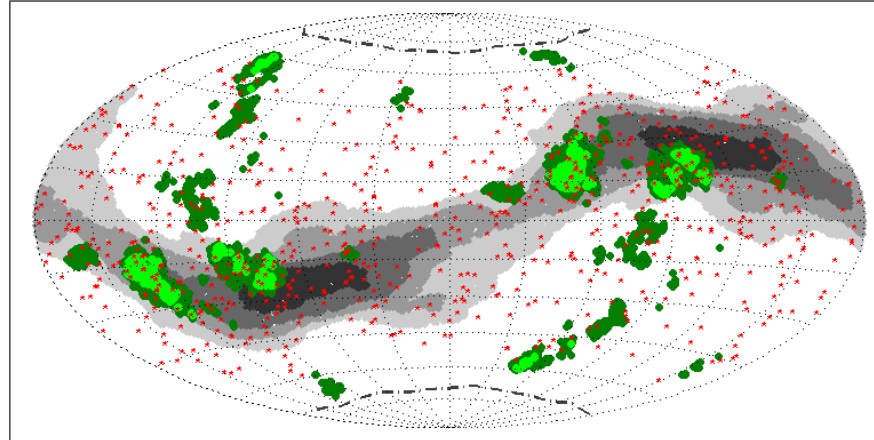


● At least 3σ deviation in number of polarisation points

● At least 2σ deviation in number of polarisation points

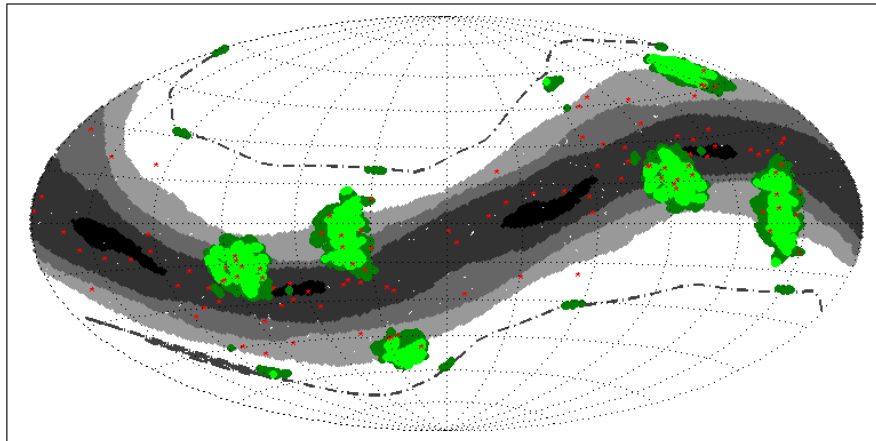
Polarisation Map Whole Sample

(355 quasars)



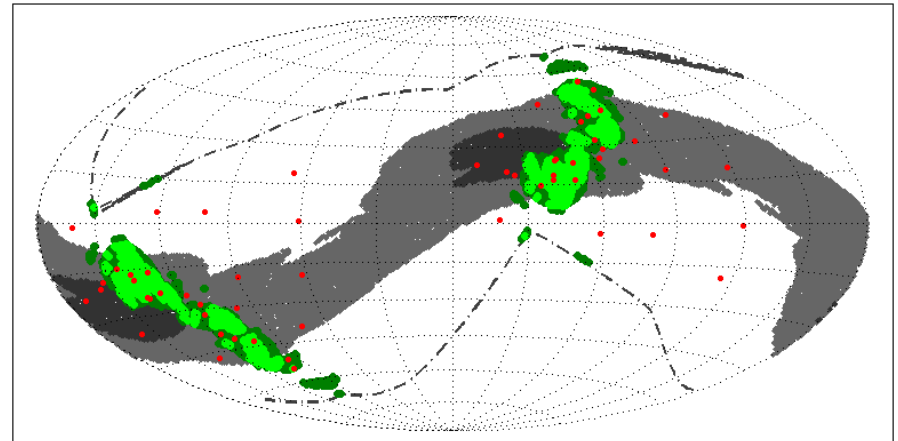
Polarisation Map A1 region

(56 quasars)



Polarisation Map A3 region

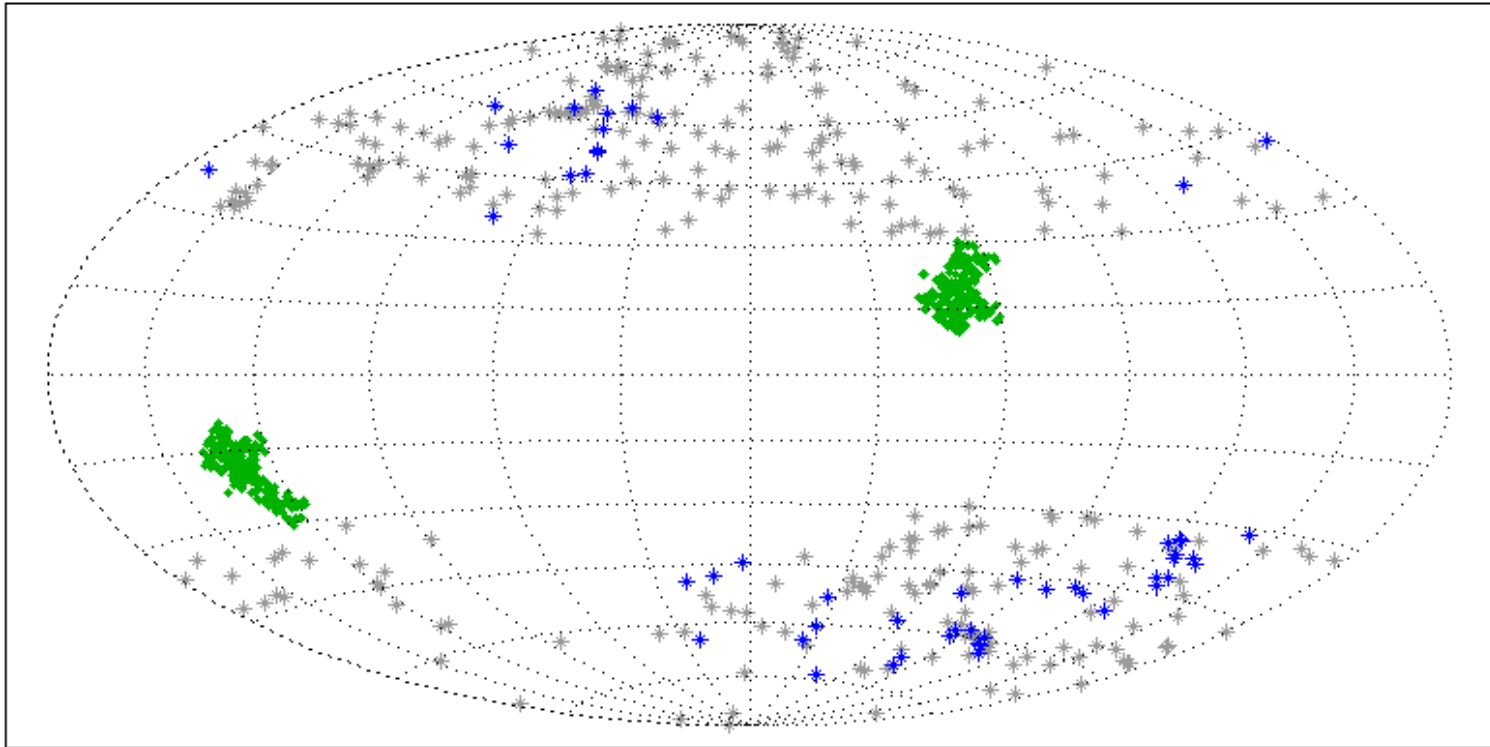
(29 quasars)



Polarisation space for the A1 and A3 regions from [\[Hutsemékers et al. 2005\]](#)

- At least 3σ deviation in number of polarisation points
- At least 2σ deviation in number of polarisation points

Sensitivity to the Antipodal Correlation



Conclusion

- a coordinate-invariant method
- confirmation of anisotropies on polarisation orientations
- objective cuts on physical parameters (RA, dec, z , p_{lin})
- preferred axis definition
- probe the geometry of the signal

DIRECT PROBE OF PHYSICS STANDING BEHIND THE
ANISOTROPIC ORIENTATIONS OF LINEAR
POLARISATIONS OF QUASAR LIGHT

Soon in [\[Pelgrims V. & Cudell J.R. \(2013\)\]](#)



Thanks for your attention

References

- Hutsemékers, D. 1998, A&A, 332,410
- Hutsemékers, D., Lamy, H. 2001, A&A, 367, 381
- Hutsemékers et al., 2005, A&A, 441, 915
- Jain, P., Narain, G., Sarala, S. 2004, MNRAS, 347, 394
- Cabanac et al., 2005, ASP Conf. Series

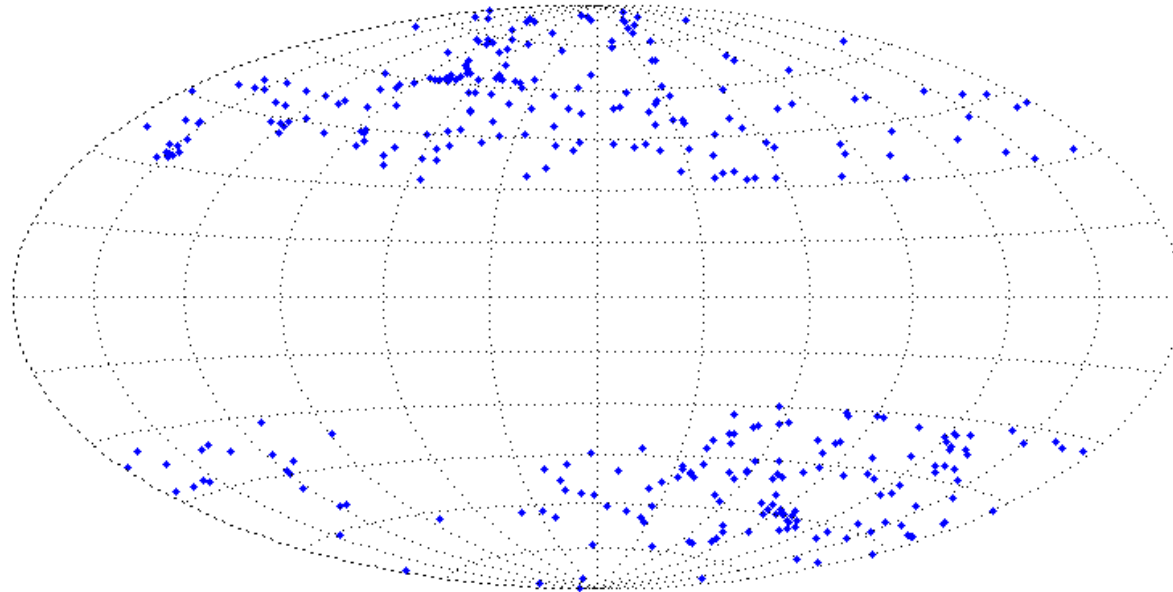
Annex

Perspectives of application this method to :

- Radio Sources
- CMB Polarisation (Planck)
- ...

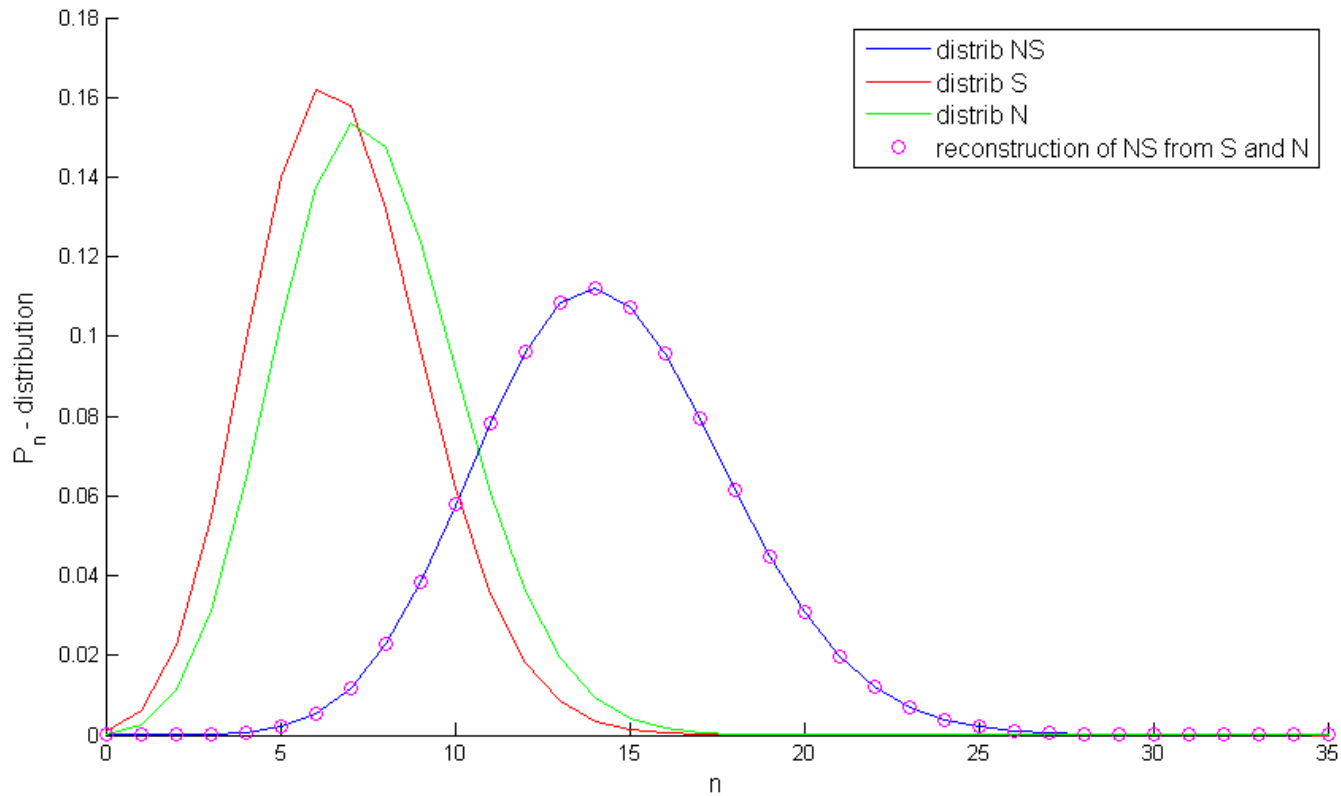
Annex

The whole sample of 355 quasars



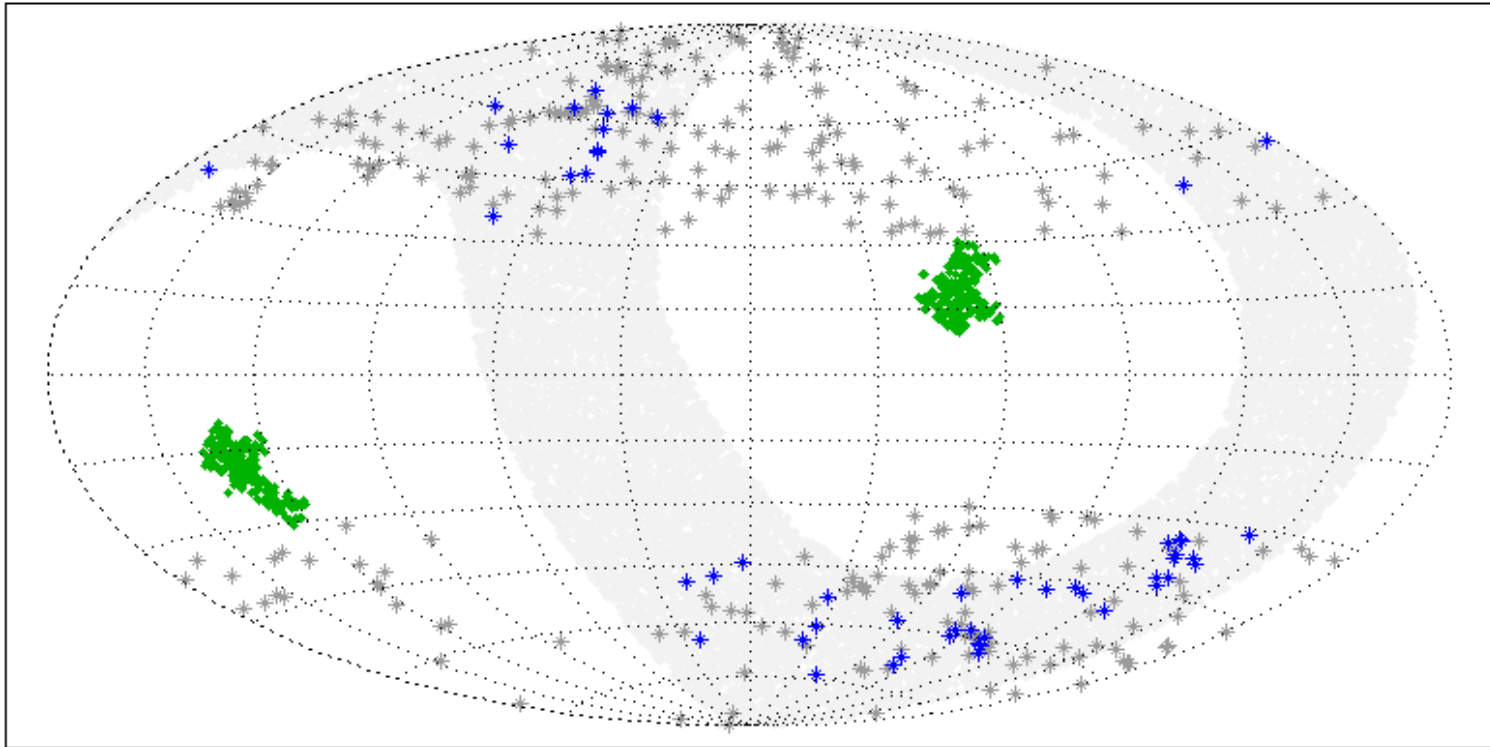
Annex

Convolution North-South



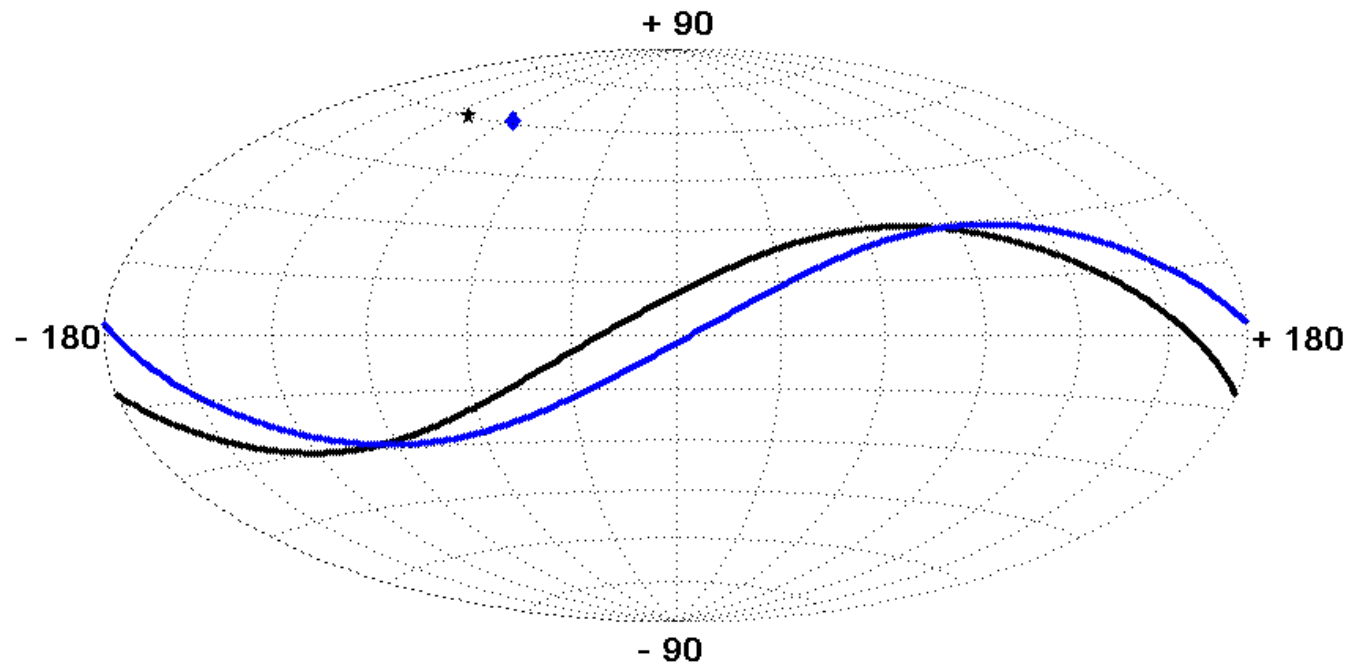
Annex

Correlation North-South and sensitive girdle



Annex

Geometrical locii, the case of 2 quasars



Annex

Individual probabilities, the case of 2 quasars and one cone

